1	Wavelet-based optical flow
2	for two-component wind field estimation
3	from single aerosol lidar data
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ABSTRACT

A motion estimation algorithm is applied to image sequences produced by 9 a horizontally-scanning elastic backscatter lidar. The algorithm, a wavelet-10 based optical flow estimator named *Typhoon*, produces dense two-component 11 vector flow fields that correspond to the apparent motion of microscale aerosol 12 features. To validate the efficacy of this approach for the remote measure-13 ment of wind fields in the lower atmosphere, an experiment was conducted 14 in Chico, California, in 2013 and 2014. The flow fields, estimated every 15 17 s, were compared with measurements from an independent Doppler li-16 dar. Time-series of wind speed and direction, statistical assessment of the 17 10-min averages and examples of wind fields are presented. The comparison 18 of 10-min averages at 100 m AGL reveals excellent correlations between esti-19 mates from the *Typhoon* algorithm and measurements from the Doppler lidar. 20 Power spectra and spectral transfer functions are computed to characterize the 21 filtering effects of the algorithm in the spatial domain. 22

23 1. Introduction

Motion estimation is a branch in the field of computer vision that develops algorithms to deter-24 mine the apparent movement of objects in sequences of digital images. Since the seminal paper 25 by Horn and Schunck (1981), the applications of these numerical methods have become numer-26 ous; they play key roles in the success of many modern technologies including bioinformatics, 27 video compression and machine vision. These techniques are also commonly found in experimen-28 tal fluid dynamics, applied for example to particle image velocimetry (PIV) (Adrian 2005). In 29 contrast to in-situ measurements which are inherently restricted to a single point of space, motion 30 estimation methods are non-intrusive and provide fields or volumes of velocity vectors and thus 31 offer a broader perspective of the flow. 32

Because of the abundance of images in the atmospheric and oceanic sciences, motion estimation 33 has been practiced since before the digital age. For example, determination of the movement of 34 cloud or water vapor features in satellite images was done prior to the work of Horn and Schunck 35 (1981) through a block-matching approach (Leese et al. 1971). These atmospheric motion vectors 36 (AMV) constitute nowadays an essential component of the observations assimilated by numerical 37 weather prediction models (García-Pereda and Borde 2014). Other modern applications involve 38 for example the recovery of glacier velocities (Scambos et al. 1992), displacements resulting from 39 landslides (Stumpf et al. 2013), surface water flows (Dugan et al. 2014) and breaking waves dy-40 namics (Melville and Matusov 2002). 41

Another application, similar to PIV and AMV, involves the estimation of 2D, 2-component wind field from the apparent motion in aerosol backscatter lidar data (Schols and Eloranta 1992). Thus far the motion estimation algorithms used in that context were variations of the cross-correlation method (Mayor et al. 2012; Hamada et al. 2015). In this paper, a more recent approach that was devised specifically for application to fluid motion is investigated. This algorithm, named *Typhoon*, is a wavelet-based optical flow estimator. It was previously validated with synthetic and
real PIV images (Dérian 2012). Here, as a first step, the validity of this wavelet-based optical flow
approach in the context of atmospheric lidar data is demonstrated.

The paper is organized as follows: Section 2 introduces the motion estimation framework for the wind measurement problem and the traditional cross-correlation algorithm. Section 3 presents the proposed *Typhoon* algorithm. The input aerosol backscatter lidar data is detailed in Section 4. Finally, in Section 5, estimated wind fields are validated by comparisons with remote measurements from a commercial Doppler lidar. Power spectra and transfer functions are calculated to show the filtering effect of the proposed approach.

56 2. Wind measurement and motion estimation

57 a. Wind measurement strategies

Air motion is represented by a three-component vector and may be defined at all points in the atmosphere. The wind is generally regarded as the vector consisting of two horizontal components. Active remote wind measurement techniques may be subdivided into Doppler and non-Doppler approaches.

Ground-based radars and lidars typically collect data in a spherical coordinate system. Doppler radars and lidars directly measure only the radial (line-of-sight) component of air motion. For a Doppler radar or lidar to measure the wind, specific scanning strategies and assumptions about the air motion over space and time must be made. Wind profiling describes the use of a remote sensor to provide a vertical profile of horizontal wind vectors at a single location above the surface of the earth. Alternatively, two Doppler radars or lidars, separated by some horizontal distance, may be used to probe an area from different angles and obtain a two-component wind field. This approach
is known as "dual-Doppler" (Stawiarski et al. 2013).

Non-Doppler approaches estimate wind fields from the spatial and temporal movement of features observed by the instrument. Eloranta et al. (1975) provided some of the first remote wind measurements by lidar in the lower atmosphere. Since that time, hardware and software has advanced greatly and a small number of validation experiments have been conducted, e.g. Mayor et al. (2012). Meanwhile, other fields, in particular experimental fluid dynamics, have developed similar approaches to retrieve motions. This concept is also known to the computer vision community, where it is associated with the wide family of *motion estimation* techniques.

b. Fluid motion estimation: the vision approach

The idea of using the *apparent motion of tracers* to infer the invisible underlying fluid flow is 78 not new. It "could probably be traced far back in history to the first time a person possessing 79 the concept of velocity watched small debris moving on the surface of a flowing stream" (Adrian 80 2005). Many visualization methods have been developed, such as using droplets, dye, smoke or 81 shadows for the purpose of revealing fluid flow structures and dynamics (Van Dyke 1982). This 82 led in particular to the well-known PIV techniques, which have been used in experimental fluid 83 dynamics for almost 30 years (Adrian 2005). Our 2D, 2-component wind measurement approach 84 fits in the motion estimation context: the tracers are the aerosol *features*, visualized by the lidar 85 system, and the motion estimation technique is usually the cross-correlation. This configuration 86 is very comparable to PIV, with the important differences that the distribution of aerosols in the 87 atmosphere (the "seeding" of the flow) cannot be controlled, and that the images are not of indi-88 vidual particles, but instead of a field that approximately represents particle concentration (Held 89 et al. 2012). In these aspects, this problem is closer to AMV computation. An important difference 90

is the temporal and spatial resolutions covered by these two approaches: typically, on the order of
15 seconds and 10 meters for the considered lidar data, versus 15 minutes and several kilometers
for geostationary satellite imagery (García-Pereda and Borde 2014).

⁹⁴ c. Motion estimation framework

Motion estimation aims to recover the apparent displacements within a sequence of images. The time and space variations of an observable image quantity are used to infer the underlying motion field occurring in the image plane between two consecutive frames of the sequence. In this work, input images are the scans provided by the lidar, the movements of the variations of aerosol backscatter intensity are used to estimate the wind field.

In the following, the scan domain is noted as $\Omega \subset \mathbb{R}^2$. The observable backscatter intensity is noted as $I_n(\mathbf{x})$ at pixel $\mathbf{x} = (x_1, x_2) \in \Omega$ and at discrete time $t_n, n \in \mathbb{N}$. The *apparent displacement* between two consecutive scans I_n, I_{n+1} is a 2D vector field **u**:

$$\mathbf{u}(\mathbf{x},t_n) = \begin{pmatrix} u_1(\mathbf{x},t_n) \\ u_2(\mathbf{x},t_n) \end{pmatrix}$$

¹⁰³ This displacement is measured in pixel units and occurs over the time $\delta t_n = t_{n+1} - t_n$ s. If the scan ¹⁰⁴ has a resolution of δx m pixel⁻¹, an estimation of the *instantaneous wind velocity* **v** in m s⁻¹ is ¹⁰⁵ therefore given by:

$$\mathbf{v}(\mathbf{x},t_n) = \frac{\delta x}{\delta t_n} \mathbf{u}(\mathbf{x},t_n). \tag{1}$$

As such, the motion is assumed to be stationary during the time step δ_t .

¹⁰⁷ Velocity components v_1 , v_2 are the *in-plane* components, that is, they belong to the image plane. ¹⁰⁸ Due to the very low value of the elevation angle of the lidar scan plane (typically < 6°), these ¹⁰⁹ components coincide with the horizontal wind components (usually denoted *u*, *v* in atmospheric sciences). The *out-of plane* component (normal to the scan plane), which remains unestimated, thus corresponds to the vertical component *w*.

The question of the accuracy of motion estimation techniques is often raised. The answer is 112 complex, since it involves the data characteristics (spatial, temporal resolutions), the information 113 given by the visualization method (the image content), and the underlying motion field itself. In 114 the current context, the later contributions are difficult to quantify, as they depend largely on the 115 conditions (e.g., the presence of particulate matter, the scales and variability of the wind field). 116 However, assuming ideal conditions and a perfect model, errors related to the resolution of data 117 may be quantified. If displacements are measured as integers on the image grid, the systematic 118 error is ± 0.5 pixel, which then gives $\pm 0.5 \delta_x / \delta_t$ m s⁻¹ for each motion component. In practice, 119 various interpolation techniques allow for sub-pixel estimation, reducing this error. The error can 120 be also lowered by using a smaller δ_x and/or a larger δ_t . However, for a given motion field, a 121 smaller δ_x results in larger apparent displacements, which can be more challenging for estima-122 tion algorithms. On the opposite, larger δ_t leads to less accurate perception of the instantaneous 123 velocity, since the assumption of stationarity of the motion field is less valid over longer periods. 124

Any motion estimation technique features two main aspects. The first one, known as the *data* 125 *model*, describes the link between observations I (the aerosol backscatter intensity) and the under-126 lying unknown displacement u. This model should take into account the nature of observed data 127 and its relevant dynamics. Then, as an inverse problem, motion estimation is usually ill-posed. 128 The second aspect is therefore the *regularization*, which is required in order to close the estima-129 tion problem. The regularization may also provides information where the data model fails locally. 130 The various estimation techniques feature different data models, regularizations or implementation 131 strategies. 132

¹³³ *d. The cross-correlation algorithm, concept and limitations*

The cross-correlation technique performs independent, local motion estimations on subregions (blocks) of the scan domain. It consists in correlating a block of the first scan I_n with a translated block of the second scan I_{n+1} ; the translation vector **u** which induces a correlation peak is considered to be the displacement at the center of the block (Schols and Eloranta 1992). The estimation problem, presented in its basic form, is written as:

$$\forall \mathbf{x} \in \Omega_C, \mathbf{u}(\mathbf{x}, t_n) = \arg \max_{\mathbf{u}} \sum_{\mathbf{y} \in B(\mathbf{x})} \frac{\left[I_{n+1}(\mathbf{y} + \mathbf{u}) - \mu_{n+1}(\mathbf{x} + \mathbf{u})\right] \left[I_n(\mathbf{y}) - \mu_n(\mathbf{x})\right]}{\sigma_{n+1}^2(\mathbf{x} + \mathbf{u})\sigma_n^2(\mathbf{x})},$$
(2)

where $\Omega_C \subset \Omega$ is the set of block centers (and therefore the set of locations of estimated vectors), $B(\mathbf{x})$ is the block centered on \mathbf{x} , $\mu_p(\mathbf{x})$ and $\sigma_p(\mathbf{x})$ are the mean and standard deviation, respectively, of backscatter intensity I_p over block $B(\mathbf{x})$. Note that in practice, this cross-correlation function (CCF) is computed using the FFT for computational efficiency.

In this case, the data model is the CCF (2) itself; the regularization is implicitly given by the 143 size of block $B(\mathbf{x})$ which should be large enough to contain reliable information, yet as small as 144 possible to resolve small scale motions. Typically, neighboring blocks overlap by 50%, so that the 145 estimated motion field is *sparse* (fewer motion vectors than pixels). Each vector is the result of 146 a single *independent* problem, which makes the CCF algorithm pleasingly parallel (Mauzey et al. 147 2012). This cross-correlation approach and its numerous variants have become widely used in PIV 148 (Adrian and Westerweel 2010); in geosciences it is often applied to satellite imagery to retrieve for 149 instance glacier velocities (Scambos et al. 1992). It is also the standard method to derive AMVs 150 (Schmetz et al. 1993; García-Pereda and Borde 2014), and has given good results with aerosol 151 backscatter lidar data, as shown in Schols and Eloranta (1992), Mayor and Eloranta (2001) and 152 Mayor et al. (2012). 153

However, this method as presented in (2) is not exempt from drawbacks. First, the displace-154 ment within an entire block $B(\mathbf{x})$ is explained by a single vector $\mathbf{u}(\mathbf{x})$, which implies that this 155 displacement is assumed to be uniform (constant) over the block. The larger the block, the less 156 likely this assumption is to be true. Yet, as overly small blocks may result in uncertainties due to 157 lack of information, "large" blocks are usually preferred. This leads to the second point: as dis-158 placements occurring within large blocks are likely not uniform, the estimated $\mathbf{u}(\mathbf{x})$ corresponds 159 to a power-weighted average of the apparent displacements within the corresponding block $B(\mathbf{x})$ 160 (Hamada 2014), which results in an over-smoothed motion field. To address these issues, this 161 study proposes to evaluate a recently developed motion estimation algorithm dedicated to fluid 162 flows. 163

3. Typhoon algorithm

Early attempts with a different class of motion estimation methods, often called optical flow, 165 were conducted in 2010 on the CHATS¹ dataset and led to promising results (Dérian et al. 2010). 166 Since then the authors developed a new version of the algorithm based on a wavelet framework, 167 named Typhoon. The extensive description of the algorithm is largely mathematical and details 168 regarding the design of the data-model and the regularization can be found in Dérian et al. (2013) 169 and Kadri Harouna et al. (2013), respectively. In the following, an overview of the method and the 170 improvements made to achieve real-time wind estimation from aerosol backscatter lidar imagery 171 are provided. 172

¹Canopy Horizontal Array Turbulence Study, near Dixon, CA, 2007 – see Patton et al. (2011).

a. Optical flow, from observations to motion

The proposed approach has two major differences with respect to the cross-correlation algorithm presented above. First, this wavelet-based optical flow uses a *global* formulation: all vectors $\mathbf{u}(\mathbf{x})$ of the displacement field \mathbf{u} are estimated simultaneously by solving a single problem, whereas the cross-correlation approach in (2) has as many independent problems as vectors $\mathbf{u}(\mathbf{x})$. Second, this method provides a *dense* estimate, that is to say one displacement vector at every point \mathbf{x} of the scan domain Ω , whereas the CCF solution is usually sparse. The estimate is obtained by minimizing a functional, similar to an energy, defined over the whole scan domain:

$$\mathbf{u} = \arg\min_{\mathbf{u}} \left\{ \frac{1}{2} \int_{\Omega} [f_{data}(I, \mathbf{u})]^2 d\mathbf{x} + \frac{\alpha}{2} \int_{\Omega} [f_{reg}(\mathbf{u})]^2 d\mathbf{x} \right\}.$$
(3)

 f_{data} is the data model that depends on observations *I* and unknown displacement **u**, while the regularization f_{reg} depends on **u** only. The parameter $\alpha > 0$ balances the two terms and is fixed by the user.

¹⁸⁴ The data model used in *Typhoon* is known as the displaced frame difference (DFD):

$$I_{n+1}(\mathbf{x} + \mathbf{u}(\mathbf{x}, t_n)) = I_n(\mathbf{x}).$$
(4)

It is analogous to finding the displacement field **u** that "warps" an image into the next one. This model assumes the *consistency of backscatter intensity* along the trajectory of an aerosol feature during the time interval $[t_n; t_{n+1}]$, that is to say an aerosol feature will present the same intensity, the same "signature", in both scans I_n , I_{n+1} . Therefore any phenomena inducing a significant change in intensity, such as turbulent diffusion or out-of-plane motion, can possibly lead to *false apparent motions*.² Such phenomena are not uncommon, but it can be reasonably assumed that the time scales at which they act are significantly larger than the inter-scan time-step δt_n , so that

²False apparent motions refer here to illusory motions of aerosol features that do not correspond to the horizontal wind.

the DFD (4) remains valid. It is also important to note that from formulation (3), the data model is not *strictly* enforced. Instead, the solution achieves a balance between trying to follow the model on one hand and the regularization on the other – hence the role of the parameter α , which allows the user to give more weight to one term over the other.

Regularization schemes usually encourage the estimate **u** to follow some *smoothness assumption*. This work uses the most simple first-order regularization, originally introduced in Horn and Schunck (1981), which penalizes strong velocity gradients. For each displacement component u_i , i = 1, 2:

$$f_{reg}(u_i) = |\nabla u_i| = \sqrt{\left(\frac{\partial u_i}{\partial x_1}\right)^2 + \left(\frac{\partial u_i}{\partial x_2}\right)^2}.$$
(5)

Note that the square root is later cancelled by the square in (3). If the regularization is given much 200 more weight than the data model ($\alpha \rightarrow \infty$ in (3)), the solution that minimizes (3) moves toward 201 a uniform motion field (with $\nabla u_i = 0$ for i = 1, 2). The regularizer also takes precedence over 202 the data model locally where the latter is inefficient, for instance within uniform regions of the 203 input images. Other regularizers are available in *Typhoon*, penalizing, for instance, the vorticity or 204 divergence of the flow, or the gradient of vorticity, divergence; some of these schemes have proven 205 to be very efficient with PIV and water vapor satellite images (Corpetti et al. 2002). However, as 206 the regularization becomes more complex, the associated computational costs increase, which may 207 reduce the ability to achieve real-time estimation. Moreover, in the context of aerosol backscatter 208 lidar images, little to no improvement brought by the use of these advanced schemes was found. 209 This could be linked to the specificities of this lidar data, which will be detailed further in Section 4. 210

The DFD model (4) and the Horn and Schunck regularizer (5) inserted into (3) complete the motion estimation problem:

$$\mathbf{u}(t_n) = \arg\min_{\mathbf{u}} \left\{ \frac{1}{2} \int_{\Omega} \left[I_{n+1}(\mathbf{x} + \mathbf{u}(\mathbf{x}, t_n)) - I_n(\mathbf{x}) \right]^2 d\mathbf{x} + \frac{\alpha}{2} \int_{\Omega} \sum_{i=1,2} |\nabla u_i(\mathbf{x}, t_n)|^2 d\mathbf{x} \right\}.$$
(6)

A particularity of this problem is that the DFD model (4) is not linear in **u**, so that the whole func-213 tional is not quadratic. This complicates the minimization process, as the existence of a global 214 minimum is not guaranteed. This is another role for the regularization term: it convexifies the 215 functional as $\alpha \to \infty$. But, as large α values are unmanageable, to ensure a successful mini-216 mization it is important for the solution \mathbf{u} to lie "close" to the first guess.³ This calls for the use 217 of an incremental strategy, often known as "multi-resolution": the displacement field is estimated 218 following a coarse-to-fine process, starting with coarse structures of large amplitudes, and progres-219 sively refining toward smaller scales. This last point motivates the use of the *wavelet framework*. 220

²²¹ b. Introduction to the wavelet framework

In signal processing, the spectral space is often used to analyze or exhibit some properties of a 222 given signal. The FFT leads to a representation in terms of sine and cosine functions of specific 223 frequencies. Any spatial information is lost in the process: the Fourier coefficients, which form 224 an equivalent representation of the input signal, yield no information as to where their associated 225 frequency is or is not present. This is due to the fact that the sine and cosine functions, which form 226 the basis of the spectral space, are very well localized in frequency but have an infinite support in 227 space. Conversely, looking at the signal in the physical space does not give any information on the 228 frequency content. The wavelet formalism offers a trade-off: the wavelet functions are localized 229 both in space and frequency, thus they enable access to information on the *frequency content and* 230

³which is usually the null motion field, $\mathbf{u}(\mathbf{x}) = 0 \ \forall \mathbf{x} \in \Omega$.

the spatial location simultaneously – at the cost of lower precision. A wavelet representation of a given signal consists of a *coarse approximation* of the signal, along with several sets of *details* containing spatially-localized information at various ranges of frequencies. Note that instead of frequency, the wavelet formalism prefers the equivalent but reciprocal notion of *scale*.

This multi-scale (or, multi-resolution) representation offered by the wavelet transform is the 235 main motivation to adopt wavelet bases for displacement components u_1, u_2 . It leads to a "natu-236 ral" coarse-to-fine strategy suitable to motion estimation (Dérian et al. 2013). Approximation and 237 coarse detail coefficients are estimated first, then fine-scale details are successively added until the 238 smallest scale is reached. Besides the multi-scale framework, wavelet bases also allow the rep-239 resentation of arbitrary regular functions (a 3D fluid motion field should at least be continuous). 240 While the continuity might not be a relevant assumption for the 2D field, a sufficient regularity is 241 required in order to compute the regularizing terms presented in Section 3.a, which involve spa-242 tial derivatives. Finally, these regularization schemes find a relatively simple yet very accurate 243 implementation in that context (Kadri Harouna et al. 2013). Similarly to the Fourier transform, 244 the wavelet transform is a linear, separable⁴ operator, with fast algorithms (fast wavelet transform, 245 FWT) for computational efficiency. Wavelets are also used in many fields, from signal denois-246 ing to video compression; Mallat (2008) discusses an extensive presentation of the theory and 247 applications. 248

²⁴⁹ Conceptually, the use of wavelet bases does not lead to significant changes to the estimation ²⁵⁰ problem (6). Each motion component u_i is expressed as the inverse transform (reconstruction) of ²⁵¹ its corresponding wavelet coefficients c_i :

$$u_i = W_{inv}(c_i), \quad i = 1, 2,$$

⁴The 2D transform is obtained by combining two 1D transform, first along rows then along columns.

where W_{inv} denotes the inverse wavelet transform. The set of wavelet coefficients $\{c_1, c_2\}$ thus is the unknown to the estimation problem.

254 c. Recent improvements

The original algorithm detailed in Dérian et al. (2013) would accept square images only. If input images were rectangular, they had to be padded to turn them square, which increases the computational burden. The current version has been modified to accept rectangular images.

The main improvement is the result of redesigning the code to run in "real-time". To keep up 258 with real-time, the estimate of wind field $\mathbf{v}(t_n)$ from scans I_n , I_{n+1} must be complete by the time 259 the next scan I_{n+2} is made available, with the inter-scan time-step δt_n typically on the order of 260 10 to 20 seconds. Since the whole motion field is estimated simultaneously, the number of vari-261 ables is quite large: a dense estimate from 512×512 pixel images represents about half a million 262 unknowns. Wavelet transforms lie at the core of the estimation process. Each evaluation of the 263 functional (6) requires two inverse FWTs (to reconstruct the displacement **u** from its coefficients) 264 and two forward FWTs (to compute the gradient). In order to achieve the necessary reduction in 265 computation time, the low-level functions of the algorithm – in particular, the wavelet transforms 266 - were rewritten in CUDA language, which enables it to execute on NVIDIA's graphic processing 267 units (GPU). GPUs designed for scientific computing rely on several thousands of small com-268 puting units, thus providing massive parallelization capabilities. The CUDA version of *Typhoon* 269 running on an NVIDIA GeForce GTX Titan is 10 to 100 times faster than the original version 270 (Mauzey et al. 2014), and is sufficient to meet the real-time requirements. 271

4. Application to aerosol backscatter data

The results presented hereafter have been obtained from data collected by the Raman-shifted 273 Eye-safe Aerosol Lidar (REAL) (Mayor and Spuler 2004; Spuler and Mayor 2005; Mayor et al. 274 2007; Spuler and Mayor 2005) in 2013 and 2014 in Chico, California. The REAL is a ground-275 based, scanning, elastic backscatter lidar operating at a wavelength of 1.54 microns, with a pulse 276 energy typically between 120 and 170 mJ, a pulse rate of 10 Hz and a pulse duration of 6 ns. It 277 employs 40 cm diameter optics and an analog direct detection receiver. The backscatter signal is 278 sufficiently strong from a single pulse that averaging over multiple pulses is not required. This 279 section describes the input scan data as well as the preprocessing steps. 280

281 *a. Data preprocessing*

²⁸² Before motion estimation takes place, the raw signal delivered by the REAL must be prepro-²⁸³ cessed. Lidar data is sampled on a polar grid, with the lidar at the origin. Each scan is composed ²⁸⁴ of *shots*, with a shot being a 1D array of backscatter samples, uniformly spaced along the range *r* ²⁸⁵ every 1.5 m, collected at a given angular position θ from a single laser pulse.

The raw backscatter intensity $I_{raw}(r,\theta)$, with the range *r* and the azimuth angle θ , corresponds to the actual backscatter signal $\beta(r,\theta)$ and an additive noise $\varepsilon(r,\theta)$.

$$I_{raw}(r,\theta) = \beta(r,\theta) + \varepsilon(r,\theta)$$

The noise ε combines contributions from the atmosphere and the instrument and can be modeled by a random variable which follows a normal distribution of mean μ_{θ} and standard deviation σ_{θ} . Values of μ_{θ} , σ_{θ} change slightly from one shot to another, hence their dependency in θ ; they can be estimated for each shot from background data. As explained in Mayor et al. (2012), first the ²⁹² noise mean is subtracted:

$$I_0(r,\theta) = I_{raw}(r,\theta) - \mu_{\theta} = \beta(r,\theta) + \varepsilon_0(r,\theta),$$

with $\varepsilon_0(r, \theta) = \varepsilon(r, \theta) - \mu_{\theta}$ the now centered random noise. The raw signal-to-noise ratio (SNR) is computed at that point:

$$SNR_{raw}(r,\theta) = \frac{I_0(r,\theta)}{\sigma_{\theta}}.$$
 (7)

Shots are then multiplied by the square of the range to compensate for the one-over-range-squared decay of the backscatter β :

$$I_{r^2}(r,\theta) = r^2 I_0(r,\theta) = r^2 \beta(r,\theta) + r^2 \varepsilon_0(\theta)$$

²⁹⁷ Note that the noise amplitude now increases as the square of the range. For optimal results, it is
 ²⁹⁸ then essential to discard irrelevant noisy data, which is discussed further.

After conversion to decibels, shots are filtered in the range dimension. The low-pass median filter of length 7 points (10.5 m) removes high-intensity spikes typically caused by hard-targets such as birds and insects, while the high-pass median filter of length 333 points (500 m) removes the very large structures to reveal local fluctuations. Figure 1 presents an example of preprocessed backscatter data (panel a), along with the corresponding raw SNR (7) (panel b).

304 *b.* Detecting coherent features

Two different aspects complicate the motion estimation process. First, due to the nature of backscatter data, the raw SNR (7) decays as one-over-range-squared. Typically, for the REAL operating in Chico, CA, the SNR resulting from a single laser pulse drops below 5 at r = 3 km. Such high levels of noise in the far range are challenging for optical flow. Second, for the purpose of motion estimation, a good SNR in the near range does not necessarily imply useful information. For instance, coherent features can be absent from a region of the scan, yielding much uncertainties as to the underlying wind field in that region.

In order to maximize the quality of the results, the scan areas presenting no coherent aerosol features are discarded. Because of the regularization schemes provided by optical flow (Section 3.a), wind vectors estimated over noisy areas *could* be relevant. However judging so proves to be difficult, as often even a basic visual confirmation is impossible in noisy regions. Hence, it is safer to simply discard the noisy image data before motion estimation.

To detect the presence of coherent aerosol features, the *image SNR* is used. It is defined as the ratio of the *local* standard deviation of coherent signal $\sigma_{\beta}(r, \theta)$ to the *local* standard deviation of noise σ_{ε} :

$$SNR_{img}(r,\theta) = \frac{\sigma_{\beta}(r,\theta)}{\sigma_{\varepsilon}(r,\theta)}.$$
 (8)

This ratio is estimated from the autocovariance function of preprocessed data $I(r, \theta)$. For every point (r, θ) , the autocovariance C_l is computed along the range from data in [r - l/2; r + l/2]. Then, the local variance of coherent signal is given by the average of coefficients at lag 1 and -1:

$$(\sigma_{\beta})^2 = 0.5 (C_l(-1) + C_l(1))$$
,

while the local variance of noise is obtained from the 0-lag coefficient and σ_{β} :

$$(\sigma_{\varepsilon})^2 = C_l(0) - (\sigma_{\beta})^2.$$

An example of image SNR is shown in Fig. 1 panel c. A 256-point window was used to compute the autocovariance, corresponding to l = 384 m.

From the image SNR, a valid data domain is computed for each scan. It is assumed that the best data is in the near range, therefore the valid domain is simply defined by a far-range boundary. For each shot (azimuth θ), this far-range boundary is given by the smallest range $R(\theta)$ above which the image SNR remains below a threshold τ fixed by the user:

$$\forall \theta, R(\theta) = \min_{R} \left\{ R : \forall r > R, SNR_{img}(r, \theta) < \tau \right\}.$$
(9)

Finally, a low-pass median filter of width 25 points and a Gaussian filter of parameter $\sigma = 2$ points are applied to the set of $R(\theta)$, to exclude small isolated features and smooth the boundary. An example of mask representing the valid data domain is shown in Fig. 1 panel d, using $\tau = 3$.

c. Correction of image distortions

A lidar scan does not correspond to an instantaneous view of the aerosol distribution. The shots 334 that compose the scan are acquired sequentially. In the event of high wind speeds, this leads to 335 apparent distortions of the aerosol features in the lidar images, which in turn causes the estimated 336 motion to be biased. This issue was first noted by Sasano et al. (1982) who proposed an iterative 337 correction method. Assuming that the aerosol features are transported *without deformation* by a 338 uniform wind vector, scans can be warped to reconstruct an approximated instantaneous view of 339 the aerosols, thus improving the accuracy of motion estimation. In this study, implementation 340 proceeds as follows for a given scan pair: 341

 $_{342}$ i. Estimate the displacement field **u** from the pair of scans with the *Typhoon* algorithm.

ii. Convert to velocity field \mathbf{v} using (1).

iii. Correct both scans for distortions using wind field \mathbf{v} , following Sasano et al. (1982). The time of the beam at the center of the scan is used as the reference time.

iv. Repeat i–iii until mean wind speed $|\bar{\mathbf{v}}|$ changes by either less than 1%, or less than $0.25\delta x/\delta t$. Typically, it requires 2-3 iterations.

The correction step (iii) is carried on the polar grid data. After correction, backscatter data is no longer known on a regular polar grid, but instead is at scattered locations.

350 *d. Cartesian gridding*

After preprocessing, masking and correction for distortions, the backscatter data is interpolated on a Cartesian grid of spacing $\delta x = 8$ m. It is possible to perform the motion estimation directly on the original polar grid, however, as mentioned above, the correction step destroys the regularity of the mesh. Fast interpolation on large sets of scattered data can be challenging, considering real-time requirements. In this work, a CUDA implementation of nearest-neighbor interpolation was used.

357 5. Validation

A field experiment was conducted in Chico, CA, from mid-September 2013 to mid-January 358 2014, to validate the wind fields recovered by Typhoon. A Doppler lidar (DL) was deployed to 359 provide independent wind measurements. It is a pulsed, heterodyne detection Doppler lidar com-360 mercialized by HALO Photonics under the name Stream Line (Pearson et al. 2009). This model 361 was previously certified against cup anemometer measurements (Axel and Ailt-Wiard 2014), and 362 showed very good agreement to radar wind profiler and radiosonde (Päschke et al. 2014). The DL 363 has the S/N 0811-35 and was built in November 2011. It operates at a wavelength of 1.5 microns, 364 pulse energy of 20 μ J, pulse rate of 15 kHz, and a pulse duration of 150 ns. DL data was fil-365 tered following the manufacturer's indications, keeping only points for which the minimum SNR 366 intensity > 1.01. 367

Since it is not possible to retrieve a 2D 2-component wind field using a single DL, two different configurations were investigated.

• *Temporal validation*. The DL was located at 1500 m range, 15° azimuth from the REAL and operated in vertical profiling mode. Data from this configuration enable comparisons of time-series of 2-component wind velocities at the DL location. This phase of the experiment
 was conducted during September and October 2013.

Spatial validation. The DL was located on the roof of the REAL container and operated in fixed-beam mode, staring at the center of the sector scan area swept by the REAL. This configuration enables one to compare *radial wind* velocity components *along the DL line-of-sight*. Data for this second phase of the experiment were collected in December 2013 and January 2014.

The main parameters used by both systems during these two experiments are summarized in Table 1.

381 a. Temporal validation

In this experiment, the REAL scans between -15° and 45° azimuth, with a 4° elevation, every 17 s. This places the scan at 100 m AGL at the range of the DL. The DL operates in vertical profiling mode (VAD scan), providing a profile of 2-component horizontal wind vector about every 15 s.

A typical example of aerosol motion estimation is presented in Fig. 2. It features a close-up of two motion fields estimated from three successive position plan indicator (PPI) scans. The flow is relatively uniform, and can be visually identified due to a large aerosol feature that moves toward the southeast. The DL wind vectors at 100 m AGL are displayed for comparison and show a good agreement with the *Typhoon* estimates.

³⁹¹ In this paper, an effort is made to establish the potential of the *Typhoon* algorithm when applied ³⁹² to aerosol backscatter lidar data. However, quality of the data depends upon the performance of ³⁹³ the instrument and the state of the atmosphere. Therefore, we selected the days presenting the

best potential for this validation among data collected in Chico, CA from mid-September to mid-394 November 2013, with the expectation that future advances in hardware will lead to increases in 395 data quality and availability. First, due to the local typical conditions in Chico, aerosol backscatter 396 imagery is much better for this application during the daytime than during nighttime. There-397 fore, this study was restricted to daytime only. Second, the percentage of valid backscatter data 398 (Sec. 4.b), during daytime, in a 50 m radius around the DL were computed. These values are 399 plotted against the mean wind speed measured by the DL the same day in Fig. 3. With a suffi-400 cient spatial distribution of aerosol features, dense 2-component wind fields can be delivered up to 401 several km in range. Figure 4 shows an example of such wind field on a day with high speed and 402 uniform direction, with vectors available out to 4 km range. The low-SNR area in the far-range 403 were dynamically excluded. Figure 5 presents a view of a ≈ 200 m vortex, illustrating the ability 404 of *Typhoon* to extract coherent structures at intermediate scales. 405

Three specific cases are described below: light, moderate and strong wind conditions. These days are represented by solid diamonds in Fig. 3. For each case, time-series of instantaneous and 10-min averaged wind measurements are presented. 10-min averages are the reference measures for instrument validation in the wind power industry (Bailey 2012). Then, statistics on 10-min averages for the 15 days having more than 85% valid data are presented.

The VAD scan strategy used by the DL assumes that the wind is uniform throughout the swept area (Mann et al. 2009, 2010; Sathe et al. 2011; Sathe and Mann 2012); in this case this region is a disc of about 100 m radius, represented by a turquoise circle in Figs. 4 and 5. In order to compare results of the study to the DL measures, instantaneous *Typhoon* estimates are averaged in space over a similar sized area centered on the DL location.

⁴¹⁶ Occasionally, the estimation may fail and result in obvious outliers. Those outliers can be de-⁴¹⁷ tected and removed under the assumption of temporal coherence of the wind field. The *normalized* ⁴¹⁸ *median test*, commonly used in PIV (Adrian and Westerweel 2010), was implemented. Similar ⁴¹⁹ concepts are used with radar wind profilers (Weber et al. 1993). Within each 10-min window, the ⁴²⁰ median wind vector $\mathbf{v_m}$ is computed, as well as the residuals $r(\mathbf{v}) = |\mathbf{v_m} - \mathbf{v}|$ for each vector \mathbf{v} of ⁴²¹ the window. Vectors for which the residual $r(\mathbf{v})$ is twice larger than the median of residuals r_m are ⁴²² discarded.

423 1) LIGHT WIND CASE

Figure 6 shows wind speed and direction measured by the DL at 100 m AGL and estimated 424 by Typhoon for a 12-hour period starting on October 23 at 15:00 UTC. It is a light wind episode 425 with speeds remaining below 3 m s^{-1} and variable direction. Estimates are missing over a pe-426 riod approximatively covering 15:00 to 17:00 UTC. This is due to the coherent feature detection 427 presented in Sec. 4.b: no significant features were present in the region of interest at that time, 428 therefore no motion estimates are available. Then, between 17:00 and 18:00 UTC, Typhoon speed 429 and direction estimates are in systematic error. Visual inspection of the aerosol imagery reveals 430 the mixed layer growing with the entrainment zone passing through the altitude of the intercom-431 parison. It appears that the plumes and wind shear in the entrainment zone result in false apparent 432 motions that bias the motion estimations. Later, two reversals of wind direction occurred at 22:30 433 and 23:30 UTC that correspond to the passage of a vortex of diameter ≈ 200 m over the region 434 of interest (see also Fig. 5 for a spatial visualization). This microscale circulation resembles those 435 that have resulted from large eddy simulation of convective boundary layers (Schmidt and Schu-436 mann 1989; Kanak 2005; Sullivan and Patton 2011). Correlation coefficients R^2 for the 10-min 437 averaged wind components are 0.951 and 0.600 for u and v, respectively. Excluding the 17:00 – 438 18:00 UTC period with false apparent motions, R^2 values increase to 0.966 and 0.866. 439

440 2) MODERATE WIND CASE

Figure 7 shows wind speed and direction measured by the DL at 100 m AGL and estimated by *Typhoon* for a 12-hour period starting on September 17 at 15:00 UTC. This wind episode features speeds ranging 0 to 10 m s⁻¹ and direction mostly stationary except for a 2-hour fluctuating episode (corresponding to the lowest wind speeds). Wind speed is underestimated at two occasions, both corresponding to rapid and large changes in direction around 22:30 and 23:00 UTC. Otherwise, both series of data are in very good agreement. This is confirmed by the 10-min averaged wind components: correlation coefficients R^2 are 0.979 and 0.991 for *u* and *v*, respectively.

448 3) STRONG WIND CASE

Figure 8 shows wind speed and direction measured by the DL at 100 m AGL and estimated by *Typhoon* for a 12-hour period starting on October 9 at 15:00 UTC. It is a strong wind episode with speeds up to 16 m s⁻¹ and very consistent flow from the northwest direction. Both timeseries are again in very good agreement. Correlation coefficients R^2 for the 10-min averaged wind components are 0.984 and 0.929 for *u* and *v*, respectively.

454 4) OVERALL CONSIDERATIONS

Scatter plots of 10-min averaged wind components measured during the daytime⁵ for the 15 "best" days (Fig. 3) are presented in Fig. 9. They show an overall excellent agreement of *Typhoon* estimates with DL measurements at 100 m AGL: correlation coefficients R^2 are 0.995 and 0.997 for *u* and *v*, respectively. Detailed statistics on *u* and *v* are available in Tables 2 and 3. In terms of wind speed, a linear regression gives a slope of 1.000 with an offset of -0.10 m s⁻¹, R^2 coefficient

⁵"Daytime" is arbitrarily considered to be 15:00 – 01:00 UTC (10 hours).

⁴⁶⁰ is 0.991. Regarding the wind direction, the offset is 1.1° and R^2 coefficient is 0.944.⁶ This \approx 1° ⁴⁶¹ offset observed for the direction corresponds to the precision at which the DL was oriented during ⁴⁶² its deployment. The root mean square error (RMSE) between *Typhoon*'s estimates and the DL ⁴⁶³ observations is 0.29 m s⁻¹ on both on *u* and *v* components. This is slightly higher than the expected ⁴⁶⁴ systematic error of $0.5\delta_x/\delta_t \approx 0.24$ m s⁻¹ which assumes perfect data and model (Sec. 2.c). The ⁴⁶⁵ few remaining outliers mostly correspond to false apparent motions, typically occurring at the ⁴⁶⁶ beginning and end of the day as the boundary layer depth evolves.

From the time-series shown in Figs. 6, 7 and 8, it appears the variability of the wind speed 467 obtained by *Typhoon* is less than that measured by the Doppler. Figure 10 is a scatter plot of 468 turbulent kinetic energy (TKE) as measured by the Doppler and Typhoon over 10-min intervals. 469 A linear regression suggests that the TKE from *Typhoon* is about 50% smaller than the Doppler's. 470 This could be linked to the fact that *Typhoon* measures apparent displacements, which are later 471 converted to velocities (Sec. 2.c). Small-scales velocity structures, either in time or space, are 472 less accurately perceived. Using a faster scan rate is likely to improve the results. Nevertheless, 473 *Typhoon* performs better than the cross-correlation technique: the optimized algorithm presented 474 in Hamada et al. (2015) recovers 39% of the TKE on the same dataset. 475

476 b. Spatial validation

⁴⁷⁷ During this phase of the experiment, the DL was colocated with the REAL. The REAL swept ⁴⁷⁸ between 15° and 75° azimuth at 2° elevation every 17 s. The DL held its beam fixed at 45° azimuth ⁴⁷⁹ and 2° elevation, measuring the radial velocity component as a function of range and time. DL ⁴⁸⁰ measurements were integrated over one second, with a range gate of 48 m. The temporal resolution

⁶When dealing with circular data such as angles, the slope for the linear regression should be fixed to 1. The offset and R^2 only are computed, see e.g. Fisher (1995).

⁴⁸¹ of DL measurements is therefore much finer than that of the REAL flow fields, and conversely for ⁴⁸² the spatial resolution (see Table 1).

Instead of holding the DL beam fixed, a PPI sweeping strategy identical to the REAL's could have been used, thus allowing the comparison of radial components over the whole scan domain. However, two arguments support the choice of a fixed beam:

- With a moving beam set-up, the integration time for DL measurements was reduced to less than 0.1 s. This would cause the SNR to decrease very rapidly. Typically in Chico the maximum range with useful data would be on the order of 1500 m, significantly below that of the REAL's.
- The radial velocity fields collected by the DL would suffer from the same distortions as the backscatter data (Section 4.c), so correcting these distortions would be challenging.

The data used for the spatial validation were recorded in December 2013 and January 2014. In 492 Chico, CA, the days are shorter and the air is cleaner during this season than in the autumn when 493 time-series data were collected. Both the DL and the REAL are affected. Data are of lower quality 494 than shown for the temporal validation. The availability of 10-min averages falls below 50% after 495 3 km for both instruments and at 5 km it is below 5%. Therefore, the analysis is restricted to the 496 first 3 km. Furthermore, it should be noted that the prevailing wind direction during this time over 497 Chico, CA is northwesterly. At 45° azimuth, the line-of-sight component corresponds mostly to 498 the cross-stream, turbulent wind perturbations. In these data, its magnitude remains mostly below 499 3 m s^{-1} . Figure 12 shows a comparison of radial velocity measured by the DL and extracted from 500 the 2-component fields obtained by Typhoon for a 8-hour period starting 8 January 2014 at 17:00 501 UTC. 502

In order to compute statistics, radial velocities were averaged. First, spatial resolution are 503 matched by averaging Typhoon velocities in space according to DL range gates, then 10-min time-504 averages are computed at every range. A scatter plot of these 10-min averages is presented in 505 Fig. 13, along with linear regression slopes, R^2 coefficient and distribution of differences. These 506 values were obtained from 8-hour periods (17:00 to 01:00 UTC) for 8 days of December 2013 and 507 January 2014. The R^2 coefficient (panel d) decreases with the range and this is expected as both 508 instruments are affected by the gradual reduction in SNR. R^2 remains above 0.95 over the first 509 1.5 km, then slowly decreases to about 0.8 at 3 km. The overall R^2 is 0.928. While the relation 510 between Typhoon and DL velocities remains linear, the slope (panel c) increases with the range, 511 from about 0.95 at 0.5 km to 1.3 at 3 km. Velocities obtained by the cross-correlation method 512 show a similar trend (Hamada et al. 2015). This leads to a theory that these discrepancies are due 513 to a mismatch in the actual elevation angles of the beams during this phase of the experiment, 514 especially considering the unbiased results of the temporal validation. At a lower elevation angle 515 and therefore lower altitude, the DL would measure lower velocities. 516

517 c. Spectral analysis

⁵¹⁸ In this section, temporal and spatial power spectra of the velocity components produced by ⁵¹⁹ *Typhoon* are presented, with the objective of characterizing the filtering effect of the algorithm – ⁵²⁰ in particular, in the spatial domain. The velocity data analyzed were collected during the daytime ⁵²¹ and within the turbulent lower atmospheric boundary layer.⁷ Therefore, an inertial subrange in the ⁵²² power spectra of the actual velocity field is expected.

⁷RHI scans collected every 15 min by the REAL during the 15 days included in the analysis show that the maximum convective boundary layer height, that typically occurred in the afternoon, ranged from 300–1200 m AGL.

The spectra are computed in natural coordinates to account for the anisotropy of atmospheric 523 boundary layer turbulence. The west-east and south-north wind velocity components are projected, 524 according to the mean wind direction, as streamwise (u_s) and cross-stream (v_n) components, such 525 that u_s carries the mean speed and v_n has a null mean. The mean wind vector is defined accordingly 526 to the investigated dimension, either in time or space. The spectra are finally averaged together 527 according to the mean wind speed, using bins of 0-4 m s⁻¹, 4-8 m s⁻¹, 8-12 m s⁻¹ and 12-528 16 m s^{-1} , in order to exibit their evolution with increasing wind speed and turbulent kinetic energy. 529 The resulting power spectral densities (S) are multiplied by frequency (f) or wavenumber squared 530 (κ^2) so that an inertial subrange would appear as a -2/3 slope and white noise would appear as 531 +1 slope. 532

533 1) TEMPORAL POWER SPECTRA

During the experiment, the REAL collected PPI scans every 17 s and one RHI scan every 15 min. 534 The RHI scan resulted in an 30 s interruption of the PPI scan sequence. The scan strategy of the 535 DL provided vertical profiles of horizontal winds every 15 ± 1 s. Since the FFT requires data 536 points at a uniform time interval, the Typhoon and DL wind measurements were interpolated to 537 a 5 s time series. From the 5 s time series data, we computed power spectra over consecutive 538 10 min intervals. The 10-min mean wind vector was used for the projection in natural coordinates 539 and the binning of spectra, as defined above. The resulting spectra have a Nyquist frequency of 540 0.029 Hz (34 s period) for the Typhoon velocities and 0.033 Hz (30 s period) for the DL. The 541 lowest frequency is 1.67×10^{-3} Hz (10 min period). 542

The spectra are presented in Fig. 14. Those from the Doppler lidar are consistently higher than the spectra from *Typhoon*, this is consistent with our observation that the TKE measured from Doppler velocities are larger than those from *Typhoon* (Fig. 10). The temporal spectra appear to ⁵⁴⁶ become flatter as the mean wind speed increases. We hypothesize that this may be caused by the ⁵⁴⁷ challenges that both *Typhoon* and the DL face under windy conditions. For the DL, increased ⁵⁴⁸ variability of the actual wind velocity field in the VAD sample area results in more error in the ⁵⁴⁹ horizontal wind vector estimate. The increased error appear as noise at these time scales and ⁵⁵⁰ flatten the spectrum. For *Typhoon*, windy conditions result in larger horizontal displacements ⁵⁵¹ between scans and faster deformation of aerosol coherent structures.

552 2) SPATIAL POWER SPECTRA

An independent observation of the 2-component 2-D velocity field does not exist for comparison with those produced by *Typhoon*. A dual-Doppler lidar set up could have provided it, but would have doubled the cost and complexity of the project. Therefore, to investigate the integrity of the vector flow fields in space, spatial power spectra are considered.

A 1 km diameter circular area is considered, centered on the DL at 1.53 km range. All of the 557 vectors within this area (from a single flow field in time) are used to compute the spatial mean wind 558 vector, which then define a natural coordinate system. Vectors of the flow field are interpolated on 559 a 128×128 point grid (1024 m \times 1024 m) that is centered on the DL and aligned with the natural 560 coordinate system, and then projected as streamwise (u_s) and cross-stream (v_n) components. This 561 operation was performed for each flow field independently and resulted in 30092 flows fields over 562 15 days. At 4° elevation, the 1024 m \times 1024 m area covers a range of altitudes from about 50 m 563 to 150 m AGL. A possible impact of this is that the turbulence statistics within this sloped domain 564 are slightly inhomogeneous. Nevertheless, for each component u_s , v_n , the 2D power spectral 565 densities computed by FFT for each flow field are averaged together according to the mean spatial 566 wind speed. Finally, slices of the resulting 2D power spectra were extracted along the streamwise 567 and cross-stream directions. This results in four 1D spectra for each wind speed bin: along the 568

streamwise and cross-stream directions, for each of the streamwise and cross-stream components. The Nyquist wavenumber is $\kappa/2\pi = 0.0625 \text{ m}^{-1}$ (16 m wavelength), the lowest wavenumber is 9.77 × 10⁻⁴ m⁻¹ (1204 m).

The spectra in the top row of Fig. 15 show the TKE increasing as expected as function of wind speed. Each spectrum has a maximum amplitude at low wavenumbers. We hypothesize that the peak corresponds to one over the Eularian length scale, and is within the energy containing range (Kaimal and Finnigan 1994). However, the spectra are steeper than $\kappa^{-2/3}$. We attribute this to two factors. First is the likely absence of aerosol features at all scales and all locations in the scan area at all times. Second is the regularization used in *Typhoon* which favors a smooth motion field, especially as the estimation reaches the smallest scales.

A transfer function describes the ratio of two spectra and, in the present work, represents the 579 attenuation of the actual wind field caused by the motion estimation as a function of wavenumber. 580 A highly idealized spectrum is constructed to serve as the reference. This is done by first locating 581 the maximum of each mean spatial spectra shown in Fig. 15. We assume that the observed power 582 at wavenumbers smaller than the peak in the spectra are accurately captured by the algorithm and 583 serve as a proper approximation of the power at those large scales. For scales smaller than the 584 peak, we extrapolate by a power-law dependence through the higher wavenumbers that mimics 585 the inertial subrange (a $\kappa^{-2/3}$ spectrum). The transfer functions are then given by the ratio of the 586 observed mean spectra over the idealized spectrum, and presented in the bottom row of Fig. 15. 587 The higher the wind speed, the more energy is missing at small scales. The ratio typically drops 588 below 50% at scales of ≈ 100 m ($\kappa/2\pi \approx 0.01$ m⁻¹) for the highest wind speeds, and ≈ 75 m for 589 the lowest. 590

591 6. Broader perspectives and conclusions

613

In a recent paper, entitled *Review of turbulence measurements using ground-based wind lidars*, 592 Sathe and Mann (2013) conclude that "Non-coherent detection may also provide possible new 593 ways to estimate atmospheric turbulence, but to our knowledge it does not, so far, challenge the 594 capabilities of coherent Doppler lidars." In this paper, we have (1) introduced a new motion 595 estimation method; (2) made the first direct comparisons of the "non-Doppler motion estimation 596 approach" with Doppler lidar; and (3) computed transfer functions to estimate the filtering effect 597 of the approach. The new motion estimation method resolves finer spatial scale flow details than 598 the traditional cross-correlation algorithm (Hamada et al. 2015). The comparisons in the time 599 domain reveal excellent correlation in terms of 10-min averages, close for example to standards 600 expected of commercial floating lidars (Carbon Trust 2013). However, the proposed approach still 601 underestimates the TKE by about 50% of what is observed by Doppler lidar. It is important to 602 keep in mind that the Doppler also provides a filtered version of the actual flow field. 603

Two horizontal components are required for wind speed and direction. The proposed approach delivers dense 2-component wind fields from a single lidar, whereas a single Doppler only produces a single component. In addition to wind resource assessment, wind fields such as delivered by *Typhoon* from REAL imagery enable the visualization and investigation of meteorological phenomena such as vortices and fronts. They also open the possibility of studies in the Lagrangian reference frame, and the tracking of flow structures or aerosol features.

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 Thomas provided helpful suggestions for data analysis.

APPENDIX A

614	Mathematical Symbols
615	• \forall for all;
616	• \subset subset of;
617	• \in in (belonging to);
618	• \mathbb{N}, \mathbb{R} the sets of natural and real numbers, respectively.
619	APPENDIX B
620	Parameters of Typhoon
621	Unless specified, results were obtained using the following parameters for Typhoon:
622	• version: cuTyphoon 1.0;
623	• wavelet basis: Daubechies, 10 vanishing moments;
624	• wavelet scales: 8 details scales considered and estimated;
625	• pyramid steps=1, scaling factor=50%;
626	• data model: DFD, smoothing kernel $\sigma = 0.5$;
627	• regularization: Horn & Schunk, $\alpha = 0.05$;
628	• data range: $[-0.5, 0.5]$, with normalization, without histogram matching.
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747		(Sec. 5.a)

	Temporal	validation	Spatial validation		
	Doppler	REAL	Doppler	REAL	
scan type	VAD	PPI	STARE	PPI	
azimuth (°)	_	[-15;45]	45	[15;75]	
elevation (°)	-	4	2	2	
range (km)	_	[0.5;5.5]	[0;5]	[0.5;5.5]	
components	2	2	1	2	
δx (m)	-	8	48	8	
δt (s)	15 ±1	17	1	17	

TABLE 1. Main parameters of DL and REAL measurements for the temporal and spatial validation experiments.

TABLE 2. RMS error, linear regression variables (slope, offset), correlation coefficient R^2 , number of points and recovery percentage w.r.t. DL reference for the 10-min averaged wind component *u* (west-east), for the temporal validation results (Sec. 5.a).

case	RMSE (m s ⁻¹)	slope	offset (m s ⁻¹)	<i>R</i> ²	# points	% recovery
light	0.17	1.047	-0.01	0.951	61	84.7
moderate	0.29	0.974	-0.05	0.979	72	100
strong	0.33	0.938	0.32	0.984	72	100
15 days	0.29	0.989	-0.03	0.995	892	99.1

TABLE 3. RMS error, linear regression variables (slope, offset), correlation coefficient R^2 , number of points and recovery percentage w.r.t. DL reference for the 10-min averaged wind component *v* (south-north), for the temporal validation results (Sec. 5.a).

case	RMSE (m s ⁻¹)	slope	offset (m s ⁻¹)	<i>R</i> ²	# points	% recovery
light	0.27	0.660	-0.02	0.600	61	84.7
moderate	0.23	0.999	0.00	0.991	72	100
strong	0.34	0.897	-0.72	0.929	72	100
15 days	0.29	1.001	0.03	0.997	892	99.1

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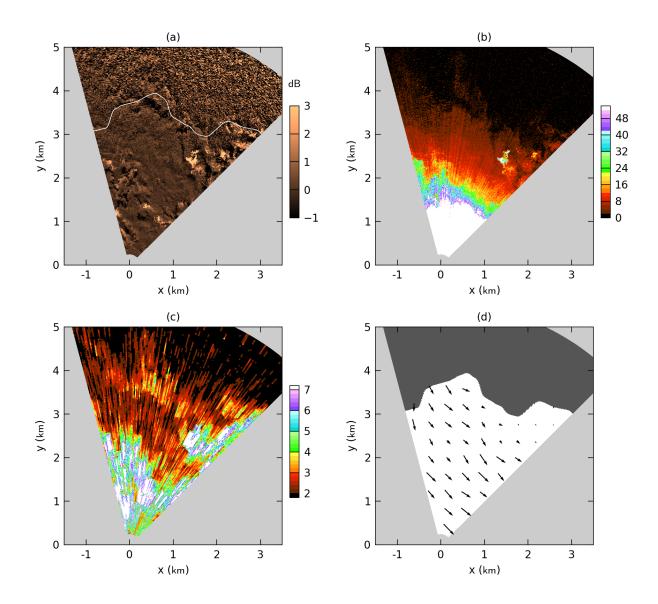


FIG. 1. Example of preprocessing applied to a horizontal scan collected on 3 October 2013 at 23:14:10 UTC. Panel (a) is the preprocessed backscatter data. Panel (b) is the raw SNR (7), revealing a $1/r^2$ decay. Panel (c) is the image SNR (8) computed using a 384 m window. Panel (d) is the valid data domain computed from image SNR. Motion is estimated in the white area only, excluding far-range noisy regions. The far-range boundary (9) of this area is also shown in (a) as a white line. Resulting, decimated vector flow field has been added to the valid area in (d).

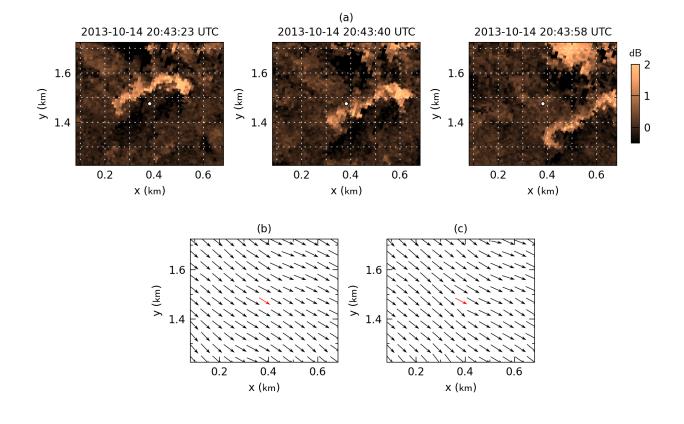


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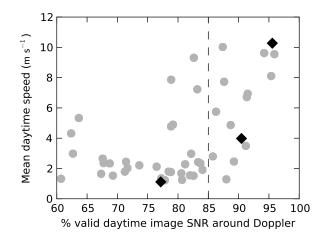


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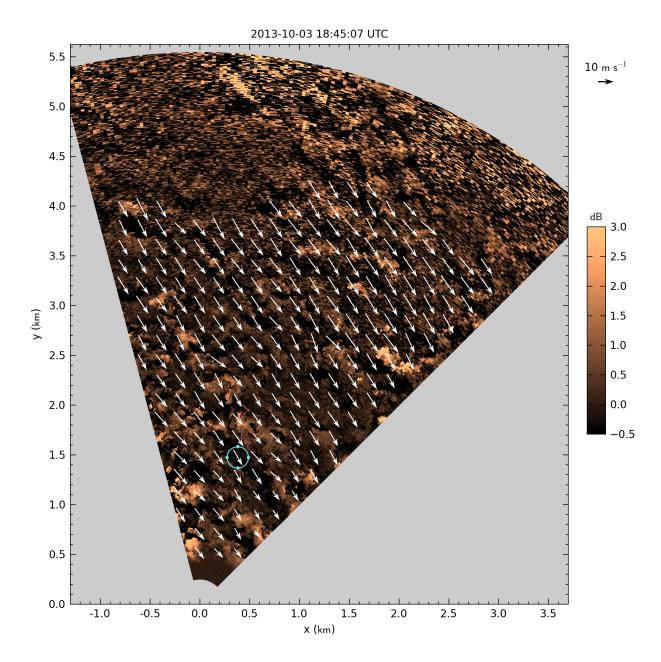


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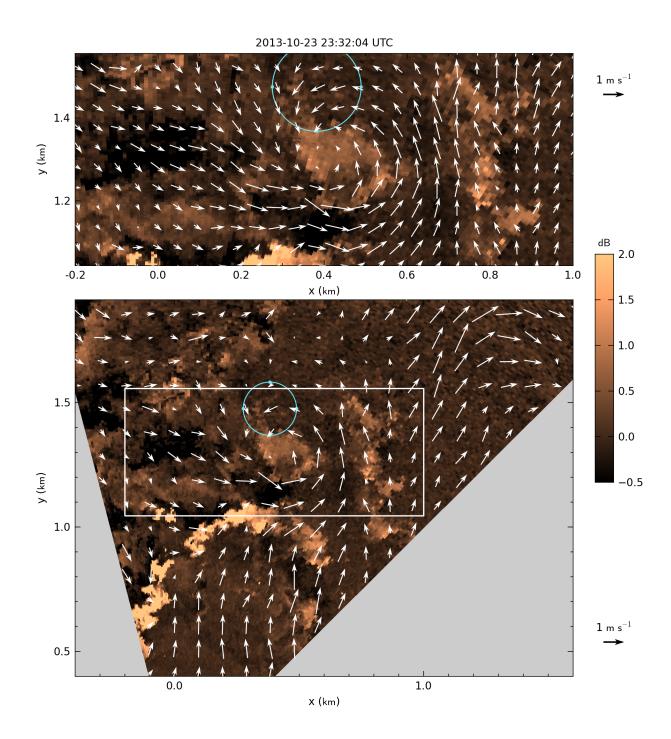


FIG. 5. Wind field obtained by *Typhoon* 23 October 2013 at 23:32:04 UTC, superimposed on the first scan of the pair used for estimation. The upper panel shows a close up on a vortex of radius \approx 200 m. The motion field was decimated along both dimensions by a factor of 6 and 12 for the top and bottom panels, respectively. The turquoise circle represents the cone section sampled by the DL during the VAD scan.

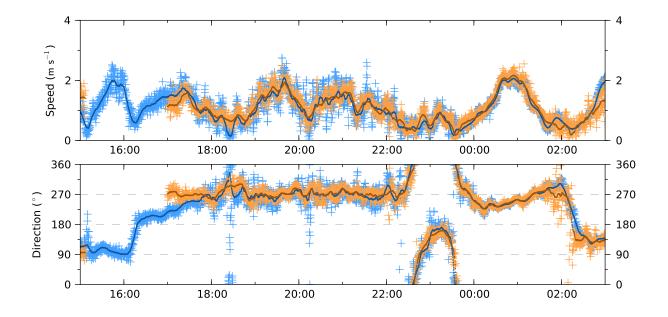


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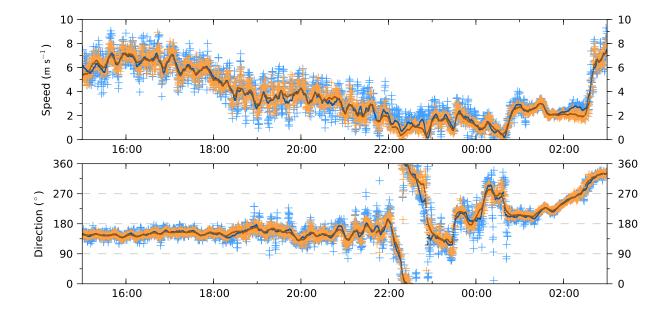


FIG. 7. Times series of wind speed (top) and direction (bottom) as measured by the DL (blue) and estimated by proposed method (orange), for a 12-hour period starting 17 September 2013 at 15:00 UTC (moderate wind case). Light + markers are instantaneous values, darker lines are the 10-min rolling averages.

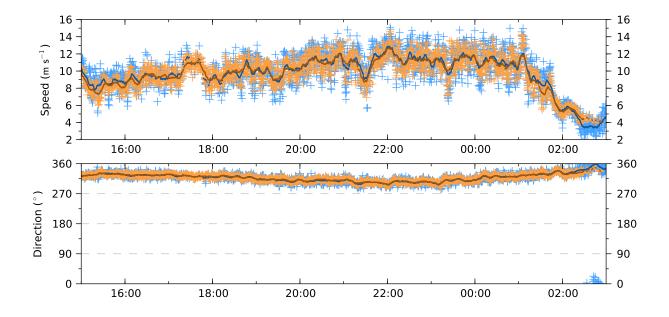


FIG. 8. Times series of wind speed (top) and direction (bottom) as measured by the DL (blue) and estimated by proposed method (orange), for a 12-hour period starting 9 October 2013 at 15:00 UTC (strong wind case). Light + markers are instantaneous values, darker lines are the 10-min rolling averages.

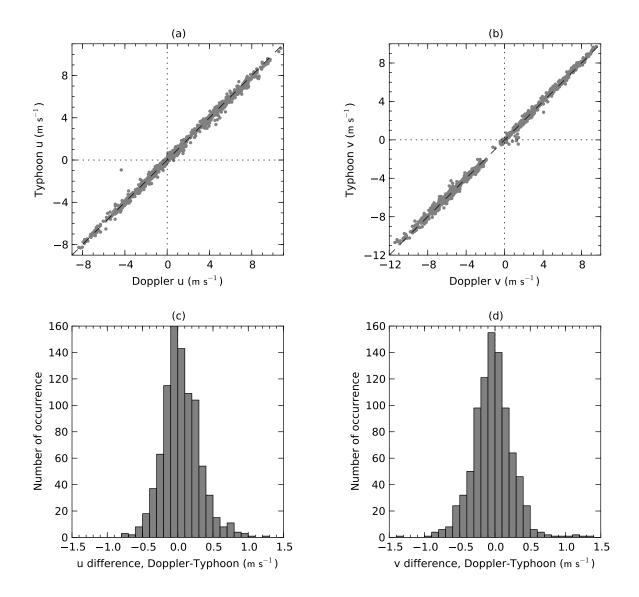


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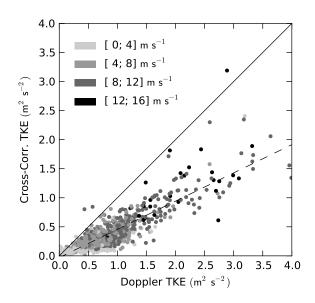


FIG. 10. Scatter plot of the TKE measured over 10-min intervals, by the DL at 100 m AGL (horizontal axis) versus estimated by the proposed method (vertical axis) – 892 points total. The gray shading indicated the mean wind speed measured over the interval. A linear regression (dashed line) gives a slope of 0.49 and offset of -0.03.

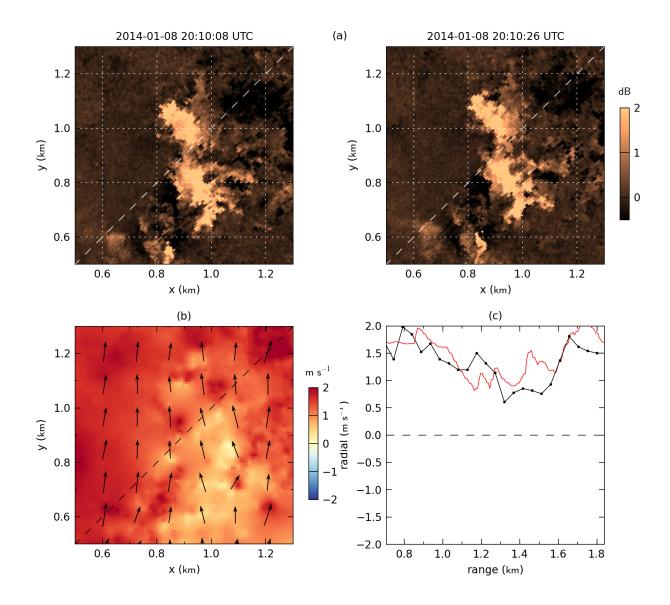


FIG. 11. Illustration of the experimental design for the spatial validation of motion estimation vectors. Panel 878 (a) shows subsets of 2 consecutive PPI scans collected on 8 January 2014 by the REAL. The displayed area is 879 a close-up centered on the DL line-of-sight at 45° azimuth (dashed white line). The copper shading indicates 880 the intensity, in dB, of aerosol backscatter. A large aerosol feature is being advected north. Panel (b) shows the 881 velocity field (black arrows) estimated by Typhoon from these two scans; the vector field was decimated by a 882 factor of 15 along both dimensions for the sake of visualization. The color shading in the background indicates 883 the corresponding radial velocity. Panel (c) compares the radial velocities measured by the Doppler (black line) 884 and extracted from the 2-component field estimated by Typhoon (red line), along the DL line-of-sight. 885

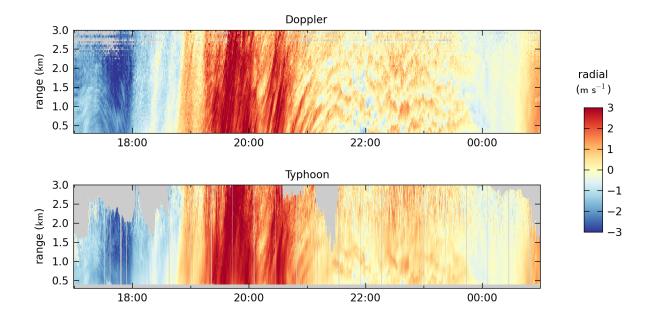


FIG. 12. Comparison of *radial wind component* at 45° azimuth and 2° elevation measured by the DL (top) and estimated by proposed method (bottom), as a function of time (horizontal axis) and range (vertical axis), for a 8-hour period starting 8 January 2014 at 17:00 UTC. Gray shading indicates missing or discarded data.

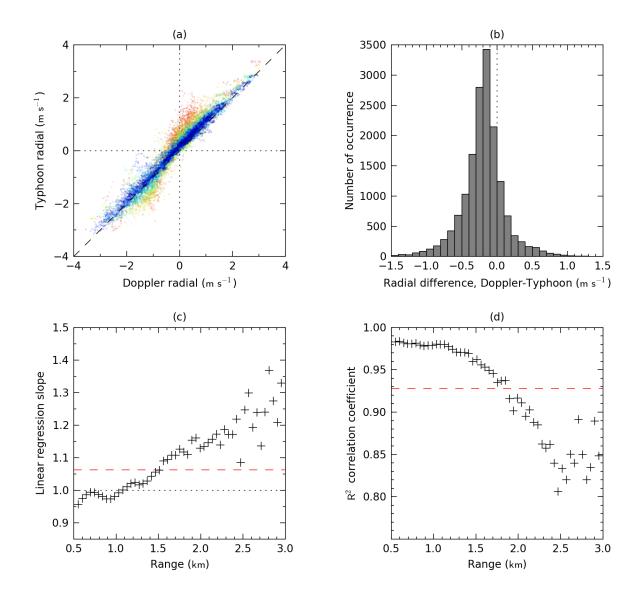


FIG. 13. Panel (a), scatter plot of 10-min averaged radial wind component measured by the DL (horizontal axis) versus estimated by the proposed method (vertical axis). Color indicates the range, from blue (0.5 km) to red (3 km). Panel (b), histogram of differences. Panel (c), slope of linear regression as a function of range. Panel (d), R^2 coefficient as a function of range. Dashed red lines indicate overall slope and R^2 values.

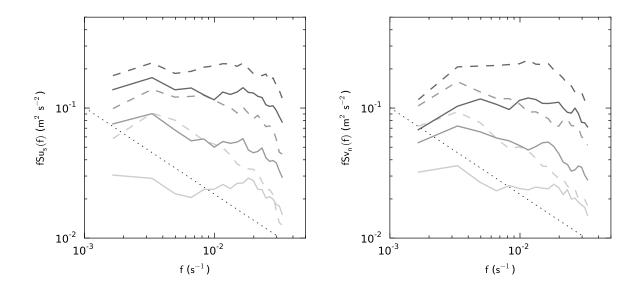


FIG. 14. Temporal spectra for stream-wise component u_s (left) and cross-stream component v_n (right) obtained by *Typhoon* (solid lines) and the DL (dashed lines). The shadings from light to dark gray correspond to wind speed ranges of [0;4], [4;8], and [8;12] m s⁻¹. The dotted line represents the -2/3 slope of the inertial subrange predicted by theory.

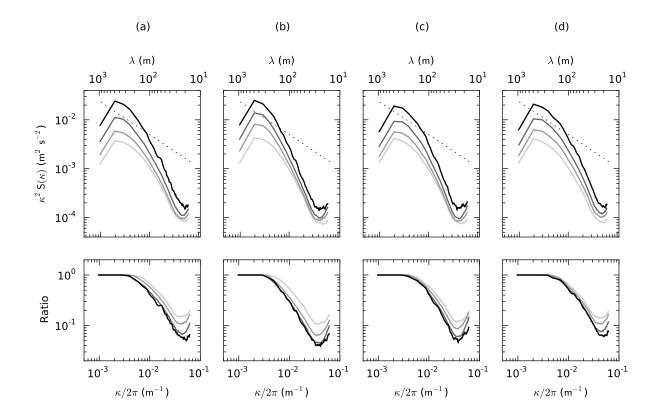


FIG. 15. Slices of 2D power spectral density (top) and corresponding transfer functions (bottom), for streamwise component u in the streamwise (a) and cross-stream (b) directions, and cross-wise component v in the streamwise (c) and cross-stream (d) directions. The shadings from light gray to black correspond to wind speed ranges of [0;4], [4;8], [8;12] and [12;16] m s⁻¹. The dotted line represents the -2/3 slope of the inertial subrange predicted by theory.